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Re-brightening of Galactic Center X-ray Transient AX J1745.6-2901

ATel #16106; *Mark Reynolds (Ohio State/Michigan), Rudy Wijnands (Amsterdam), Nathalie Degenaar (Amsterdam), Jon Miller (Michigan), Jamie Kennea (Penn State)* on behalf of a larger collaboration.

on 29 Jun 2023; 18:44 UT

Credential Certification: *Mark Reynolds (markrey@umich.edu)*

Subjects: Radio, Infra-Red, X-ray, Neutron Star, Transient

Referred to by ATel #: [16436](#), [16642](#)

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We report on our ongoing Neil Gehrels Swift Observatory monitoring observations of the Galactic center (Degenaar et al. 2015). In a recent observation (obsid: 00096991050), a transient X-ray source is detected at a position consistent with that of the known eclipsing and bursting X-ray transient AX J1745.6-2901 (atel #[14788](#), #[13453](#), #[13150](#), #[10323](#), #[9551](#)). This observation was taken on MJD 60122.76669 (Tstart: 2023 June 28 @18:23UT) and had a duration of ~0.8ks. This source lies approximately 1.5' southwest of Sgr A*.

A spectrum is extracted from a circular region ($r=18''$) centered on this source and background from an annular region $30''-50''$ around the source position. In order to characterize this source, spectral fits were carried out in Xspec, utilizing the c-statistic. We measure a net 2.0 - 10.0 keV count rate of 0.12 ± 0.01 ct/s. All quoted uncertainties are at the 1-sigma confidence level.

Assuming a powerlaw continuum, we measure:

$$N_H = (21.7 \pm 10.5) \times 10^{22} \text{ cm}^{-2}$$

$$\Gamma = 1.04 \pm 1.0$$

$$\text{Flux} = (3.4 \pm 0.1 - 1.6) \times 10^{-11} \text{ erg/s/cm}^2 \text{ (2.0-10.0 keV, observed)}$$

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cstat/dof = 67.6/85

Assuming a blackbody continuum, we measure:

$$N_H = (13.8 \pm 6.5) \times 10^{22} \text{ cm}^{-2}$$

$$kT_{\text{bbrad}} = (2.6 +1.7 -0.7) \text{ keV}$$

$$\text{Flux} = (3.6 +0.2 -1.7) \times 10^{-11} \text{ erg/s/cm}^2 \text{ (2.0-10.0 keV, observed)}$$

cstat/dof = 69.5/85

At a distance of 8 kpc, the observed luminosity ($L_x \sim 2 \times 10^{35} \text{ erg/s}$) indicates a new active phase for this transient. We note that there is evidence for the source in our previous observations on 2023 June 08/07 at an observed flux of $\sim 2 \times 10^{-12} \text{ erg/s/cm}^2$ (2.0-10.0 keV, for an assumed powerlaw with $\gamma=1.8$). AX J1745.6-2901 was not detected on 2023 June 6th or in the weeks prior ($f_x < 1 \times 10^{-12} \text{ erg/s/cm}^2$ -- 2.0-10.0 keV).

The Swift/XRT monitoring campaign website can be found at: <http://www.swift-sgra.com>

References:

Degenaar et al., 2015, JHEAp, 7, 137

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