Soft, absorbed X-ray spectra of the new transient IGR J17451-3022

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Soft, absorbed X-ray spectra of the new transient IGR J17451-3022

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We report Swift/XRT followup observations of the new transient IGR J17451-3022 discovered in JEM-X observations performed Aug. 22-24 (Atel #6451). Swift observed with XRT in PC mode on Sept. 5 (1648 seconds; position reported in Atel #6451), Sept. 9 (1950 s, starting at UT 10:00), and on Sept. 10 (856 s, starting at UT 2:12). IGR J17451-3022 is clearly detected in each observation, with moderate pileup.

We extracted source spectra (excluding the central 8-10" to avoid the piled-up region) and background spectra from the 3 XRT PC mode observations, constructed effective area files, and used the appropriate response matrices. We grouped the data to 50 counts/bin in the first two observations, and 25 counts in the third observation, and fit the spectra simultaneously in XSPEC with absorbed (TBABS, using Wilms et al. (2000, ApJ, 542, 914) abundances) power-law, blackbody (bbbodyrad), or bremsstrahlung models. An absorbed power-law model gave a poor fit (reduced chi-squared of 1.34 for 59 degrees of freedom), with photon indices greater than 3.3 (best-fit around 3.7). Bremsstrahlung models gave a better fit (reduced chi-squared of 1.16 for 59 degrees of freedom), with photon indices greater than 3.3 (best-fit at 6*10^{-22}). In this case, the temperature varies (from 0.77±0.04, to 0.88±0.05, then to 0.83±0.06), but the variations are not significant. The 2-10 keV unabsorbed flux increases from 1.2(±0.1)*10^{-10} ergs cm^{-2} s^{-1} to 1.9(±0.2)*10^{-10} ergs cm^{-2} s^{-1}, then to 1.6(±0.2)*10^{-10} ergs cm^{-2} s^{-1}. We attempted blackbody fitting with the absorption tied (best-fit at 6*10^{-22}). In this case, the temperature varies (from 0.77±0.04, to 0.88±0.05, then to 0.83±0.06), but the fit is markedly poorer (reduced chi-squared of 1.25 for 61 dof).

The nature of IGR J17451-3022 remains unclear. The high absorption is consistent with a black hole in the hard state. This might be explained if the system is nearly edge-on, reducing our visibility of the inner Galactic Center, the luminosity (2-10 keV Lx=1*10^{36} erg/s) would be unusually low for a soft state. This space for free for your conference.

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disk, which could also explain the varying N_H. Alternatively, the system may be located well beyond the Galactic Center; a distance of 20 kpc would bring the Lx up to $10^{37}$ ergs/s, reasonable for the soft state. The apparently varying N_H could in this case be due to variations in a more complex spectrum (e.g. a disk blackbody, plus comptonization, and perhaps another blackbody from a neutron star surface) that our limited spectra cannot reasonably constrain. Alternatively, IGR J17451-3022 may be a transient magnetar, in which case we should detect pulsations (in the 2-12 second range) in data of sufficient time resolution.

We thank the Swift team for their rapid scheduling of these observations.

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