A rainbow of cold atoms caused by a stochastic process

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Stochastic rainbow caustic observed with cold atoms


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We report the direct observation of a type of rainbow caustic. In contrast to known examples, this caustic originates from a dissipative, stochastic process. We have observed this when using cold $^{87}$Rb atoms bouncing inelastically on an evanescent-wave atom mirror. The caustic appears as a sharp peak at the lower edge of an asymmetric velocity distribution of the bouncing atoms. The stochastic process is a spontaneous Raman transition due to photon scattering during the bounce. The results are in good agreement with a classical calculation.

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Caustics are ubiquitous phenomena in nature. Some examples are the cusp-shaped patterns of light reflection on the inside of a coffee cup and the patterns of bright lines observed on the bottom of a swimming pool [1]. The prime example of a caustic is the common rainbow, which can be understood in a ray-optics picture by considering how the scattering angle of a light ray depends on its impact parameter on a water droplet [2]. Whereas the incident rays have smoothly distributed impact parameters, the outgoing rays pile up where the scattering angle has a local extremum. Such a divergence of the ray density, the caustic, appears at the rainbow angle. In atomic [3] and nuclear [4] scattering experiments, analogous rainbow phenomena have also been observed.

The examples of caustics known so far have in common that the outgoing parameter (scattering angle) is a deterministic function of the incoming parameter (impact parameter). In this paper we report on our observation of a new type of rainbow caustic which exists by virtue of a stochastic process, and distributes a single-valued “impact parameter” over a range of “scattering angles.” To our knowledge, such stochastic caustics have not been observed before.

We have observed this caustic in the vertical velocity distribution of cold atoms, after bouncing inelastically off an evanescent-wave (EW) mirror [5–8]. Previous experiments on the transverse velocity distribution of atoms bouncing elastically on corrugated mirrors [9,10] also allow an interpretation in terms of caustics. However, those were of the usual deterministic kind, where the outgoing transverse velocity is a deterministic function of the position where the atom hits the mirror. The caustic then originates from atoms reflecting from inflection points on the mirror surface.

Other experiments on inelastically bumping atoms have concentrated on the cooling properties. However, the appearance of the velocity caustic seems to have escaped attention, although the underlying phenomenon was present in time-of-flight signals in Ref. [6]. The incident atoms in our experiment are nearly monochromatic, i.e., they have a narrow velocity distribution. During the bounce the atoms are optically pumped to a different hyperfine ground state, by a spontaneous Raman transition. They leave the surface with less kinetic energy, due to the difference in optical potential. The outgoing atoms have a broad velocity distribution. This is due to the stochastic nature of the spontaneous Raman process, so that atoms make the transition at different depths in the evanescent wave. Our measurements and analysis show that the resulting velocity distribution is highly asymmetric. The caustic appears as a sharp peak at the minimum possible velocity. This divergence demonstrates the strong preference for making the Raman transition at the turning point.

The experiment is performed in a rubidium vapor cell. We trap about $10^7$ atoms of $^{87}$Rb in a magneto-optical trap (MOT) and subsequently cool them in optical molasses to 15 $\mu$K corresponding to a rms velocity spread of $\sigma_v = 3.8$ cm/s. The MOT has a $1/e^2$ radius of 0.6 mm and is centered 5.8 mm above the horizontal surface of a right-angle BK7 prism. When the falling cloud arrives at the surface, the mean velocity is 34 cm/s, and the velocity spread is $\sigma_v = 0.9$ cm/s. Note that this distribution is for fixed $z = 0$, not for fixed time. At the surface, we create an EW by a Gaussian-shaped laser beam of 0.8 mm $1/e^2$ radius, which undergoes total internal reflection.

The blue detuned EW induces a repulsive optical dipole potential $U_F(z) = U_F(0) \exp(-2\kappa z)$, where the subscript $F = 1,2$ denotes the hyperfine ground state, $\kappa = k_0/\sqrt{n^2 \sin^2 \theta - 1}$, with $k_0 = 2 \pi/\lambda$ the free space wave vector of the light, $n = 1.51$ the refractive index, and $\theta$ the angle of incidence [11–13]. The maximum potential $U_F(0) \approx I_0/\delta_F$, where $I_0$ is the incident intensity at the glass surface, and $\delta_F$ is the detuning. Here we define the detuning $\delta_{1,2}$ relative to the transition $5S_{1/2}(F = 1) \rightarrow 5P_{3/2}(F' = F + 1)$ of the $D_2$ line (780 nm), see Fig. 1.

An atom entering the EW in the $F = 1$ state is slowed down by the potential $U_1(z)$. Spontaneous Raman scattering can transfer the atom to the higher hyperfine ground state ($F = 2$). When transferred into this state, the potential acting on the atom is $U_2(z)$, which is weaker due to the increase of the detuning by approximately the hyperfine splitting $\delta_2 = \delta_1 + \delta_{15S}$. (Fig. 1). The ratio of the two potentials, $\beta = U_2(0)/U_1(0)$, quantifies the reduction in potential energy. As a result the atoms will bounce inelastically, i.e., to a lower height than their initial release position.

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Experimental data of bouncing atoms are taken by absorption imaging. After a variable time delay the atomic cloud is exposed to 20 \( \mu s \) of probe light, resonant with the \( ^5S_{1/2}(F = 2) \rightarrow ^5P_{3/2}(F' = 3) \) transition. Thus only atoms that have been transferred to \( F = 2 \) contribute to the signal. The atomic cloud is imaged on a digital frame-transfer close-coupled device camera. The presented images show an area of 2.2 \( \times \) 4.5 mm\(^2\) with a pixel resolution of 15 \( \mu m \) (Fig. 2). The initial position of the MOT is outside the field of view. The horizontal line at the bottom shows the prism surface.

A typical series with 2 ms time increments is shown in Fig. 2. Each density profile has been converted into a horizontal line sum. The solid curve is the result of a calculation described below. The amplitude of the experimental curves is rescaled such that the maximum optical density of the experimental curve coincides with the theoretical maximum value. This is the only fit parameter.

As expected, the atoms bounce up lower than their initial height. Furthermore, the spatial distribution shows another striking feature: it displays a high density peak at low \( z \), and a long low-density tail extending upward. Note also that there is a time focus, the density peak is sharpest when the slowest atoms reach their upper turning point. This density peak is an immediate consequence of a caustic appearing at the minimum possible velocity.

The caustic can be understood by considering the atoms as point particles moving in the EW potential. Consider an atom entering the evanescent wave with asymptotic velocity \( v_1 \). Its trajectory through phase space is determined by energy conservation: 
\[
U_1(0) \exp(-2\kappa z) + \frac{1}{2}mv^2 = \frac{1}{2}mu^2,
\]
and is depicted in Fig. 3 by the thick line. While it is slowed down by the EW it may scatter a photon at a velocity \( v_2 \), and be transferred to the \( F = 2 \) state. The atom continues on a new trajectory determined by \( U_2(z) \). To illustrate the formation of the caustic, possible trajectories starting at various positions in phase space are depicted as thin black \((v_p < 0)\) and gray \((v_p > 0)\) curves. For asymptotically large \( z \) the density of curves represents the outgoing velocity distribution, showing the caustic where the trajectories pileup. This distribution is similar to a rainbow. The density of outgoing trajectories diverges at the ‘‘rainbow velocity.’’ Below this velocity the intensity is zero, similar to Alexander’s dark band in the optical rainbow [2]. Above the rainbow velocity the intensity distribution decreases smoothly.

Despite the similarities, there is a crucial difference between a rainbow created by sunlight refracted by water drop-
other deterministic caustic. For example, the optical rainbow angle is determined by the index of refraction \( n \) of water. However, the energy of the photons is not changed, and, therefore, \( n \) cannot be compared to \( \beta \). The change in the position of the caustic with \( \beta \) can clearly be seen in Figs. 3(b) and 3(c). Here we measured the height of the upper turning point, as a fraction of the initial MOT height, for three values of the detuning.

In order to quantitatively analyze our experimental data, we write the optical hyperfine-pumping rate during the bounce as \( \Gamma'(z) = (1 - q)\Gamma U(z)/h\delta_i \), where \( q \) is the branching ratio to \( F=1 \). We define \( \eta(v) \) as the survival probability for the atom to reach the velocity \( v \) without undergoing optical pumping. This function decreases monotonically as \( \eta = -\Gamma'\eta \), with \( \eta(v_f) = 1 \). For \( |v| < |v_f| \), the solution is \( \eta(v) = \exp[-(v-v_f)/v_c] \), where \( v_c = 2\kappa h \delta_i/(1-q)\Gamma \), with \( m \) the atomic mass. When the pumping process takes place at a certain velocity \( v = v_p \), the atom leaves the surface with velocity \( v_f = \sqrt{v_p^2(1-\beta) + \beta v_p^2} \) (Fig. 1). This results in a distribution of bouncing velocities, resulting from atoms that were pumped while moving toward \( (v_p^-) \) or away from \( (v_p^+) \) the surface:

\[
n(v_f) = \left[ \eta(v_p^-(v_f)) \left| \frac{\partial v_p^-}{\partial v_f} \right| + \eta(v_p^+(v_f)) \left| \frac{\partial v_p^+}{\partial v_f} \right| \right] / v_c
\]

for \( \sqrt{\beta}|v| < |v_f| < |v_p| \), that diverges at \( v_f = \sqrt{\beta}v_p \). This velocity caustic originates from atoms which are pumped near the turning point. There are two reasons for scattering preferentially at this position. An atom spends a relatively long time at the turning point since its velocity is lowest there. In addition, the intensity of the EW, and thus the photon scattering rate, is highest at the turning point. The divergence is an artifact of the ray-optics description. It disappears due to diffraction when the atoms are treated as matter waves.

To compare the experimentally obtained spatial distributions with the model, we first calculated the one-dimensional phase-space density \( \rho(z,v) \). The spatial distribution is obtained by projecting \( \rho \) on the \( z \) axis. Initially the MOT is described by a normalized Gaussian, \( \rho(z,v) \propto \exp[-(z-z_0)^2/2\sigma_z^2]\exp[-v^2/2\sigma_v^2] \) with \( z_0 = 5.8 \text{ mm}, \sigma_z = 0.3 \text{ mm}, \) and \( \sigma_v = 3.8 \text{ cm/s} \). The cloud falls due to gravity and expands due to thermal motion. When it arrives at the EW, its velocity distribution is nearly Gaussian centered around \( v = 34 \text{ cm/s} \), with a spread of \( \sigma_v = 0.9 \text{ cm/s} \). When the cloud is reflected, its velocity distribution is changed in the way described above. We measure the spatial distribution of inelastically bounced atoms after a time-of-flight \( t \), where \( t = 0 \) is defined as the time that the center of the cloud reaches the mirror. To illustrate the agreement with the experimental data, the result for a time \( t = 7 \text{ ms} \) is drawn as the thick line in Fig. 2.

A very intriguing aspect of our experiment is the possibility to observe supernumeraries: interference of two trajectories with the same outgoing velocity, but with different pumping velocity \( v_p \). In the velocity distribution after the inelastic bounce, the high velocity tail is the sum of two contributions: from atoms that move towards and atoms that

![Image](https://example.com/image.png)

**FIG. 4.** (a) Density profiles at \( t = 14 \text{ ms} \) after the bounce for \( k = 1.96, 0.80, \) and \( 0.62 \mu \text{m} \). The curves have been offset by 0.015. The variation of \( k \) in this range has no influence on the position and shape of the spatial density profile. (b) Density profiles for various detunings: \( \delta_i = 67, 150, \) and \( 233\Gamma \), again with 0.015 offset. The curves are measured at the moment that the density peak reaches its highest point and have the same scale. (c) Measured height of the upper turning point as a fraction of the initial (MOT) height. This represents \( \beta \), the relative optical potential strength of the \( F = 1 \) and \( F = 2 \) ground states. The curve results from a calculation including the full level structure.
move away from the surface when they are pumped to $F = 2$. Interference between these two paths should result in oscillations in the velocity distribution. This is a nontrivial effect because it involves the spontaneous emission of a photon.

Given our experimental parameters, we expect a typical oscillation period of 1 cm/s. This spacing depends on $\kappa$, just as in a rainbow the supernumeraries depend on the droplet size. These oscillations are not yet resolved in our measurements, because in the EW field the magnetic substates are not degenerate, each producing a different velocity distribution.

In conclusion, we have observed a new type of caustic due to the stochastic distribution of a monochromatic input.

The caustic appears as a sharp peak at the lower edge of the velocity distribution of rubidium atoms, bouncing inelastically on an evanescent-wave atom mirror. The caustic bears some resemblance to the common optical rainbow. However, its dissipative character makes it uncomparable with any deterministic caustics and qualifies it as a new physical phenomenon.

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