Continuing outburst from H 1658-298

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Published in:
The astronomer's telegram

Link to publication

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Continuing outburst from H 1658-298

ATEL #8699; A. Bahramian, C. O. Heinke (Alberta), R. Wijnands (Amsterdam), N. Degenaar (Cambridge) on 15 Feb 2016; 23:21 UT

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Subjects: X-ray, Neutron Star, Transient

H 1658-298 is an eclipsing transient neutron star X-ray binary with a history of long (>~years) outbursts. MAXI detected a new outburst from this source on August 21st, 2015 (ATEL #7943). We monitored this source with Swift/XRT until it entered Swift Sun constraint on October 24th, 2015 (ATEL #8046). We have now resumed our monitoring of this source since it left the Sun constraint on January 24th, 2016.

We observed H 1658-298 on Jan. 28th, Feb 2nd and Feb. 11th with Swift/XRT (for 170, 970 and 590 seconds respectively), with the first observation in PC mode and the rest in WT mode. H 1658-298 is known to show dips, flares and eclipses, however investigating the lightcurves from these observations, we found no evidence of rapid variability on a timescale of seconds/minutes.

We extracted a spectrum from each observation and performed spectral fitting in XSpec. Fitting the spectra with absorbed powerlaw and absorbed disk blackbody models, we found that the absorbed powerlaw model provides a better fit, with a reduced chi-squared of 0.8, compared to 1.6 for a disk blackbody fit, for 89 d.o.f.

H 1658-298 showed an unabsorbed 0.5-10 keV flux between 1.5e-10 and 2.2e-10 erg/s/cm² on Jan 28th and Feb. 11th, with an increased flux 4.6e-10 erg/s/cm² on Feb 2nd. These variable fluxes are consistent with previous observations of the source in October 2015.

Assuming the absorbed powerlaw model, we found no evidence of variation in N_H between observations, with an average value of 5.5(+/- 0.4)e21 cm⁻². However we found suggestive evidence of spectral hardening on Feb. 2nd (when the flux had increased by a factor of ~2). While observations performed on Jan. 28th and Feb. 11th showed a consistent photon index of 2.3(+/-0.1), the observation on Feb. 2nd showed a photon index of 1.9 (+/-0.1).

Further Swift/XRT observations of this source have been planned. We thank the Swift team for scheduling our observations.

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