ABSTRACT

According to the prevailing cosmological model, 85% of the universe’s mass consists of “dark matter” – an elusive substance that is distributed throughout the universe and is undetectable by optical means. As of yet, the nature of the dark matter has been unknown for over 40 years. This thesis is about the history of the dark matter problem; it probes the historical conditions under which dark matter became a fundamental part of the canon of cosmology.

The dissertation traces the history of two phenomena that are currently regarded to be evidence for dark matter’s existence: the rapid rotation of galaxies, and the dynamics of galaxies in clusters. Unlike many popular historical accounts have argued, it is shown that these phenomena, as they were observed during the 1960s, were not directly interpreted as evidence for dark matter. Observations of galaxies and clusters of galaxies were only shaped into evidence for dark matter within a new hybrid scientific environment that arose in the early 1970s, that of physical cosmology. Driven by new observations and postwar technological developments, many young scholars engaged in doing a physics of the universe. Within this cosmological practice, concerns about “missing mass” in the universe arose. In 1974, two groups of astronomers and physicists combined the existing observations from galactic and extra-galactic astronomy to evidence that this missing mass indeed existed: galaxies were ten times more massive than was previously thought. In this context, dark matter came to matter.
Some of the material in this thesis has previously appeared in the following publications:


COLOPHON

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